

## Blazars and INTEGRAL

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**Abstract** We refer to analysis of the ESA INTEGRAL satellite data for specific class of active galactic nuclei - blazars. These objects represent promising sources to be observed by INTEGRAL, especially during their active states. Suitable strategy for the future analysis is proposed and discussed.

**Key words:** active galactic nuclei – blazars

### 1 INTRODUCTION

Blazars represent the most extreme class of active galaxies, they are powerful and variable. They are observed in all wavelength bands – from radio through VHE gamma frequencies, with maximum spectral output and largest variability often at gamma ray energies and contribution to the extragalactic background radiation across the electromagnetic spectrum (Giommi & Colafrancesco 2006, Giommi 2006). 66 blazars were identified as sources of  $\geq 100$  MeV emission by EGRET aboard CGRO (Hartman 1999) and 6 blazars were identified as VHE gamma sources ( $\geq 350$  GeV) by Cerenkov telescopes (Krawczynski 2004). It is obvious that blazars represent suitable targets for INTEGRAL satellite (Winkler et al. 2003) especially during active states (flares).

### 2 INTEGRAL OBSERVATIONS

The INTEGRAL observations are basically divided into the following categories: (i) AO-1, 2, 3 Program (allocated pointed observations); (ii) Core Program CP (Galactic Plane Scans, Galactic Center Deep Exposure,...); and (iii) Objects inside FOV of AO-1, 2, 3 observations.

### 3 CORE PROGRAMME OBSERVATIONS

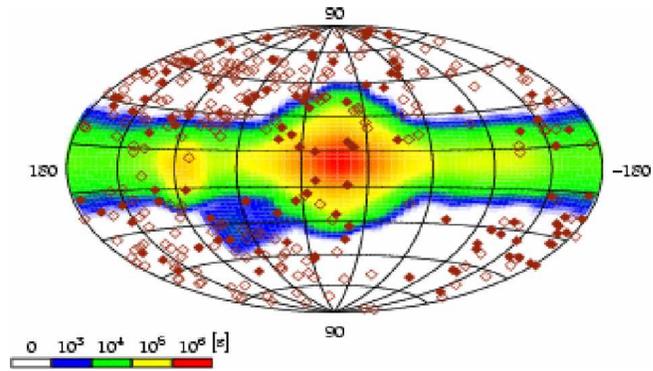
Blazars in the INTEGRAL Galactic Plane Scans (GPS) represent a promising group of objects for the study within the INTEGRAL CP. The GPS zone is usually neglected by extragalactic astronomers due to heavy obscuration: in optical,  $\sim 20\%$  of the sky is obscured by our Galaxy, while the gamma-ray telescopes on board INTEGRAL allow detection of up to few mCrabs in the most exposed GPS regions (Figs. 1, 4).

Seven optically bright (with  $V \leq 17$  mag, to be detected by the INTEGRAL OMC camera) blazars were identified within galactic scans of INTEGRAL, namely: 1ES 0647+250, PKS 0823–223 (no gamma from EGRET, grav. lensing candidate), 1ES 2344+514 (TeV gamma ray source, very close), 8C 0149+710 (BL Lac candidate?), 4C 47.08, 87GB 02109+5130 (poorly understood blazar, TeV candidate), and BL Lac (the prototype).

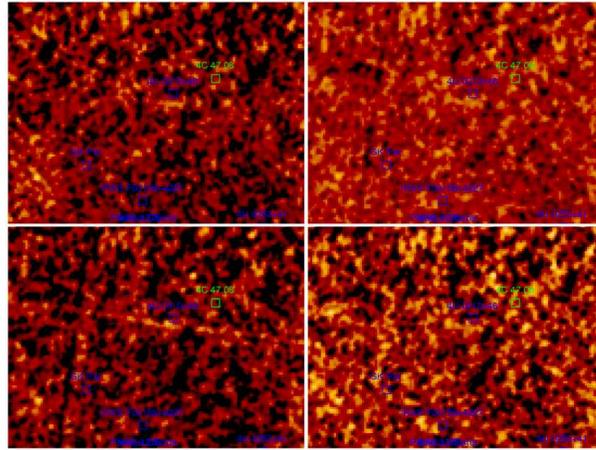
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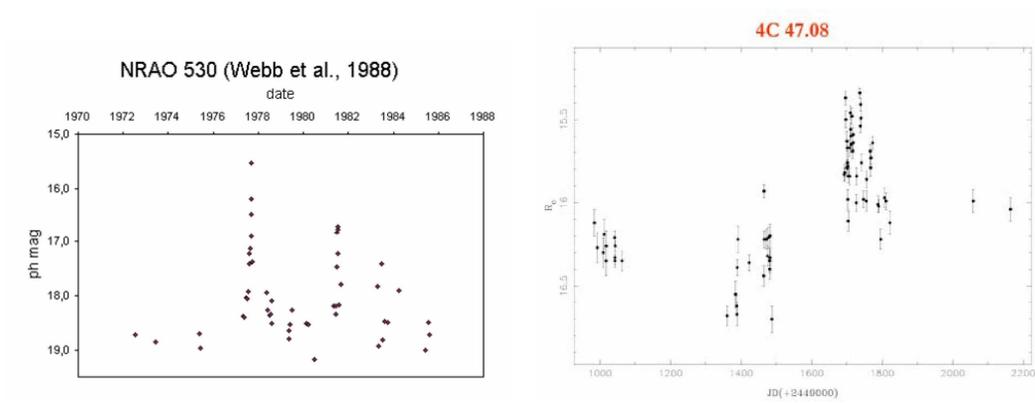
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**Fig. 1** IBIS total exposure times for blazars, CP data, revolutions 1–250.



**Fig. 2** The position of the INTEGRAL CP blazar 4C47.08 on the IBIS mosaics.



**Fig. 3** Left: Historical optical light curve of NRAO530. The source exhibits rare but large amplitude optical flares (4 mag). Right: Optical monitoring of GPS blazar 4C47.08 (Tosti et al., 2005).

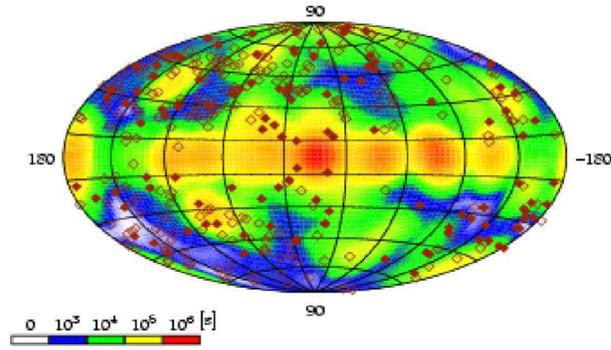


Fig. 4 IBIS total exposure times for blazars, all data, revolution 1–250.

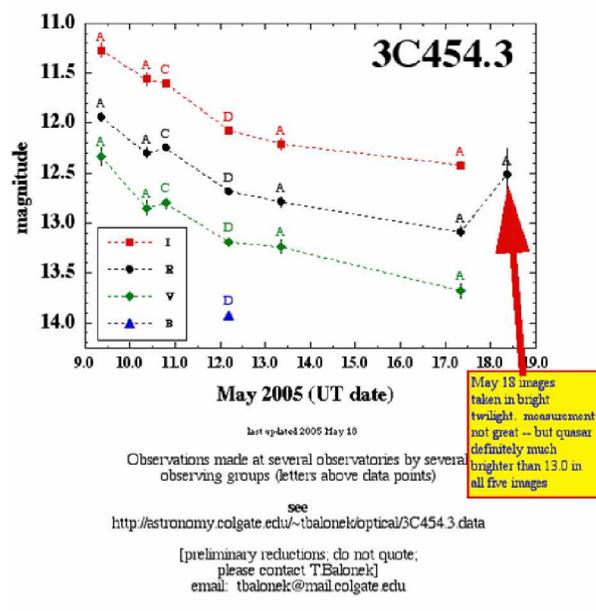


Fig. 5 The optical light curve of 3C454.3 around the time of the INTEGRAL AO-3 ToO observation.

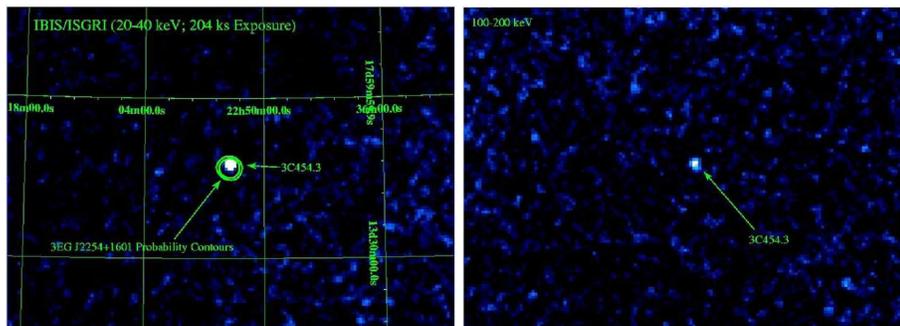


Fig. 6 INTEGRAL IBIS/ISGRI images of 3C454.3, 20–40 (left) and 100–200 keV (right).

While the prototype object BL Lac is well studied, most of the INTEGRAL GPS blazars have been poorly investigated and poorly understood so far. The study with Sonneberg Observatory Archival Plates reveals that most of these objects are optically variable, hence a gamma ray variability can be expected.

Below the detection limit of the INTEGRAL OMC on board camera is blazar NRAO530 (1730–130), which is an example of blazar with violent optical activity (4 mags within 1 month, Fig. 3). In flare, the object is expected to be much brighter also in gamma. This strengthens the role of optical monitoring and ToO program – the flare can be recognized by optical monitoring with small ( $D \sim 50$  cm) telescopes.

All the above mentioned blazars in INTEGRAL GPS have been investigated with INTEGRAL CP data (IBIS and JEM–X telescopes). We have no positive detection by high energy instruments on board INTEGRAL yet (except marginal detection of 1ES 0647+250). The targets quiet level is still below the sensitivity threshold of the instruments (Fig. 2). However, the positive detection may be possible in the future as (i) there will be more cumulative time available and (ii) the probability to see a blazar during a flare (and hence much brighter) will also increase with time.

#### 4 AO OBSERVATIONS

Additional blazars have been identified in the fields represented by the AO-1 and AO-2 observations of other scheduled targets, covered by up to 400 ksec cumulative exposure time. The analysis of these objects are in progress, as the data are becoming public.

Regarding the pointed observations of blazars by INTEGRAL, the AO-2 ToO blazar observation No. 220049 by Pian et al. (2005) has provided promising results. This collaborative proposal was based on extended optical and/or X-ray monitoring (RXTE ASM and others) of flaring activity of a large list of blazars and, alternatively, on soft gamma-ray monitoring by INTEGRAL itself (serendipitous detection of a flaring blazar in the IBIS FOV). Then ToO INTEGRAL observation was activated meeting the trigger criteria (major flaring event). The program was coordinated with the XMM Newton ToO program.

Blazar S5 0716+714 was the target of this ToO observation. This is a BL Lac object, intensively monitored at radio and optical wavelengths by Whole Earth Blazar Telescope (more than 40 telescopes, Villata et al. 2004). The ToO was triggered by optical activity - 2 outbursts up to the extreme level of  $R = 12.1$  mag (historical maximum, light increase by 1 mag in 2 weeks and 2 mags in 4 months) and, consequently, the INTEGRAL ToO observation was performed in the time interval 2004 April 2–7 (Pian et al. 2005).

The object was observed at somewhat higher (2x) gamma-ray state when in Oct 2000 (BeppoSAX ToO, Tagliaferri et al. 2003) ( $R = 12.5$  versus 12.1). There was a low signal/noise ratio and hence no intra-orbit variability study was possible, and no spectrum extraction. Within the INTEGRAL observation, the S5 0716+714 was detected only by IBIS ISGRI at 4.5 sigma, in 30–60 keV band, for a count rate of  $0.11 \text{ counts s}^{-1}$  (exposure 280 ksec). No signal above 60 keV was detected. The target was better detected at the beginning (decline), hence the useful exposure was reduced to 84 ksec. There was no positive detection in IBIS/PICsIT, SPI and JEM–X (less than 292, 6 and 6 mCrab, Pian et al. 2005).

Within the ToO observation, another extragalactic sources were observed in the IBIS FOV ( $19^\circ \times 19^\circ$  at half response) - 3 additional AGNs with higher significance than the main blazar target. These 3 AGNs were observed up to 100 keV with no intra-orbit variability study possible. However, the spectrum extraction was possible for S5 0836+710 (high  $z$  blazar of the FSRQ sub-class): single power-law spectrum, for Mkn 6 (bright Seyfert): single power law spectrum, and for Mkn 3 (bright Seyfert): broken power-law with cutoff at  $\geq 50$  keV (Pian et al. 2005).

This AO-2 blazar observation clearly confirmed that even with relatively short exposures, the INTEGRAL is an efficient tool to study bright AGNs at high Galactic latitudes and also has demonstrated the importance of high-energy instruments with large FOV and good angular resolution (like IBIS/INTEGRAL). In addition, the detection of two high  $z$  blazars (S5 0836+710 at  $z = 2.17$  and PKS 1830–21 at  $z = 2.51$ ), the most distant objects seen by INTEGRAL so far shows that INTEGRAL can also play a role in investigation of high  $z$  Universe.

Very recently, an INTEGRAL AO-3 ToO observation of 3C454.3 ( $z = 0.859$ ) was performed, with preliminary results given by L. Foschini et al. (2005, PI E. Pian with a large collaboration). This ToO was triggered by high optical (T. Balonek, VSNET alert) and X-ray (BAT Swift) activity of the source (Fig. 5). The INTEGRAL observation started 2005 May 15, at 18:40 UT, with exposure of 200 ksec. The

source was clearly detected by IBIS/ISGRI in the 20–40 keV and 40–100 keV energy bands, with a significance of 20 and 15 sigma (Fig. 6). The observed fluxes were  $1.02 \pm 0.05 \text{ cts s}^{-1} = 9.4 \pm 0.5 \text{ mCrab}$  in 20–40 keV,  $1.00 \pm 0.08 \text{ cts s}^{-1} = 13 \pm \text{ mCrab}$  in 40–100 keV, and  $1.6 \times 10^{-10} \text{ erg cm}^{-2} \text{ s}^{-1}$  in 20–100 keV band. The observed spectrum was flat, with photon index  $2.2 \pm 0.2$  and normalization  $0.13 (+0.10 - 0.06) \text{ ph cm}^{-2} \text{ s}^{-1} \text{ keV}^{-1}$ . The allocated 200 ks were however not enough to detect the object up to 400 keV, more than 400 ks would be necessary for that. It is important to continue the future INTEGRAL AO observations of blazars with longer exposures in order to fully show the importance of scientific study of blazars with this satellite.

## 5 CONCLUSIONS

For blazars in GPS, no positive detection by high-energy instruments on board INTEGRAL has been confirmed yet (except marginal detection of 1ES 0647+250 and the newly confirmed PKS 1830–211). The targets quiet level is still below the sensitivity threshold of the instruments. On the other hand, the positive detection may be possible in the future as: (i) there will be more cumulative time available, and (ii) the probability to see a blazar during a flare (and hence much brighter) will also increase with time.

For the AO blazar observation, it has been confirmed that with relatively short exposures the INTEGRAL is an efficient tool to study bright AGNs at high Galactic latitudes. This also confirms the importance of high-energy instruments with large FOV and good angular resolution. Furthermore, it is also an excellent proof that the approach by Pian et al. is the right one for variable sources and INTEGRAL-extended monitoring and ToO observations of blazars in active states. We suggest to recognize the scientific value of ToO observations of flaring blazars (and highly variable objects in general) by the INTEGRAL satellite since the observational bias mostly do not allow to see the blazars in quiescence.

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## References

- Giommi P. et al., 2006, *A&A*, 445, 843
- Giommi P., Colafrancesco S., 2004, *A&A*, 414, 7
- Foschini L. et al., 2005, *ATEL*, 497
- Hartman R. et al., 1999, *ApJSS*, 123, 79
- Krawczynski H., 2004, *New Astronomy Reviews*, 48, 367
- Pian E., Foschini L., Beckmann V. et al., 2005, *A&A*, 429, 427
- Tagliaferri G. et al., 2003, *A&A*, 400, 477
- Tosti G. et al., 2005, <http://astro.fisica.unipg.it/PGblazar/tabella2000.htm>
- Villata M. et al., 2004, *A&A*, 424, 497
- Webb J. R. et al., 1988, *AJ*, 95, 374
- Winkler C. et al., 2003, *A&A*, 411, 7